Massive Neutrinos and Flavour Violation

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Abstract

In spite of the large lepton flavour violation (LFV) observed in neutrino oscillations, within the Standard Model, we do not expect any visible LFV in the charged lepton sector ($\mu \to e, \gamma, \tau \to \mu, \gamma$, etc.). On the contrary, the presence of new physics close to the electroweak scale can enhance the amplitudes of these processes. We discuss this in general and focus on a particularly interesting case: the marriage of low-energy supersymmetry (SUSY) and seesaw mechanism for neutrino masses (SUSY seesaw). Several ideas presented in this context are reviewed both in the bottom-up and top-down approaches. We show that there exist attractive models where the rate for LFV processes can attain values to be probed in pre-LHC experiments.

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I. INTRODUCTION

Since the last couple of years of the previous century, tremendous progress has been made in our understanding of the nature of the most elusive Standard Model (SM) particles, the neutrinos. The main message of various experimental results has been that neutrinos have non-standard properties: they have masses and their flavour states mix and, indeed, with very large mixings. In the Standard Model, the phenomenon of flavour mixing is not surprising, as its presence has already been well established in the quark sector. Furthermore, the induced phenomenological implications such as K^0 - \bar{K}^0 oscillations, B_d - \bar{B}_d mixing and $b \to s, \gamma$ have been well understood as well as measured with high precision. Given this, and non-zero neutrino flavour mixing, one would expect that similar phenomenological and theoretical implications can now emanate from the leptonic sector.

The more obvious of the phenomenological implications of neutrino flavour mixing phenomena are processes like neutrino oscillations. On the other hand, the presence of this neutrino flavour mixing would also induce flavour mixing in its Isodoublet partners, the charged leptons. This mixing, induced at the 1-loop level through gauge bosons, is manifested by rare decay processes such as $\mu \to e, \gamma, \tau \to \mu, \gamma$, etc. If only the neutrinos carry this information on flavour mixing, as in the Standard Model with massive neutrinos, these processes are expected to be proportional to the ratio of masses of neutrinos over the masses of the W bosons, leading to extremely tiny branching ratios.

This situation can dramatically change if there are some new additional particles carrying lepton flavour numbers, which have very large masses and simultaneously mix among themselves. The presence of such particles could lead to enhanced branching ratios for the above processes, perhaps bringing them into the realm of observability of the present and next generations of experiments. On the other hand, non-observability of these processes can lead to strong constraints on the nature of new physics, which is expected to be present just above the electroweak scale. Such constraints already exist from the hadronic sector, albeit riddled with typical uncertainties associated with them. The leptonic flavour changing processes, on the other hand, do not suffer from these uncertainties, and thus lead to stronger constraint on any new flavour violating physics. In fact, the present constraint on BR($\mu \to e, \gamma$) has long been considered to be the most stringent constraint on any new flavour physics. To get a feeling as to where we stand, we provide here a list of present and upcoming experimental limits:

Present limits

$$BR(\mu \to e\gamma) \le 1.2 \times 10^{-11}$$
 [1]
 $BR(\tau \to \mu\gamma) \le 3.1 \times 10^{-7}$ [2]
 $BR(\tau \to e\gamma) \le 3.7 \times 10^{-7}$ [3]

Upcoming limits

$$BR(\mu \to e\gamma) \le 10^{-13} - 10^{-14}$$
 [4]
 $BR(\tau \to \mu\gamma) \le 10^{-8}$ [3]
 $BR(\tau \to e\gamma) \le 10^{-8}$ [3]

The impact of these limits could be felt in a wide class of new physics models setting in at a scale close to the electroweak scale. A particularly interesting class of models are the supersymmetric (SUSY) Standard Model(s). In these models, the supersymmetric partners of the leptons, namely the sleptons, carry the same flavour quantum numbers as the SM leptons. Since they are expected to be heavy, around the TeV scale, if flavour mixing is present in the (s)leptonic sector, large

branching ratios are expected for the afore-mentioned rare decay processes. Interestingly enough, this naturally occurs if we marry the idea of low energy supersymmetry to the mechanism of seesaw [5] giving rise to small neutrino masses (SUSY seesaw [6]). In the present review, we will mainly concentrate on this class of models.

II. SUPERSYMMETRIC MODELS AND LEPTON FLAVOUR VIOLATION

The study of flavour violation in supersymmetric models is certainly quite complicated. Indeed, to the usual intricacies involved in the FCNC computations in the SM, we add several new SUSY contributions. The main source of this difficulty is the soft supersymmetry breaking lagrangian, which can in general contain a large number of flavour violating couplings, leading to significant constraints on its parameters [7, 8]. An appealing manner to cope with these tight constraints is to consider only a particular classes of soft lagrangians, which result from models that break supersymmetry in a flavour blind manner, as in mSUGRA, Anomaly mediated supersymmetry breaking (AMSB) or its several variants, Gauge mediated supersymmetry breaking (GMSB), etc [9]. However, in general, even after choosing a particular model of supersymmetry breaking, flavour violation can still be present in the weak scale lagrangian. The various sources of flavour violation in such a case can be broadly classified as¹:

- (i). In models of supersymmetry breaking based on supergravity or superstring theories, although it is possible to achieve universality or even no-scale boundary conditions under some assumptions on the Kähler potential, non-universal soft terms are generically present in the high scale effective lagrangian [10].
- (ii). In models with flavour symmetry imposed by a Froggatt–Nielsen mechanism, flavour violating corrections to the soft potential could be potentially large [11]. More so, if the flavon fields contain SUSY breaking F-VEVs [12, 13].
- (iii). Finally, the existence of new particles at high scales with flavour violating couplings to the SM leptons (as right handed neutrinos in a seesaw model [6]) or the presence of new Yukawa interactions (as in Grand Unified Theories where quark and leptonic fields sit in the same (super)multiplets [14]) can lead to flavour violation at the weak scale. In this case, the flavour violation is communicated to the low energy fields through Renormalisation Group Equations (RGEs) [15].

Obviously, this list is not exhaustive. There can be other exotic sources which can be either additional heavy Higgs particles [16] or are related to the localisation of the fermions in higher dimensional space-time when MSSM/SUSY-GUT is embedded in a extra-dimensional model [17, 18].

In the present review, we will concentrate on the flavour violation solely due to a mechanism generating neutrino masses and mixings. To this effect, we would consider models where the strong universality is assumed at the high scale and thus point (i) would not be considered here. Similarly, we will not consider either effects generated by the imposition of a flavour symmetry as in (ii). Instead, the main aspect of this review will be to collect some salient features of the flavour violation induced by a neutrino mass model in a supersymmetric theory. As was mentioned in the introduction, the most natural and popular of them, the seesaw mechanism, will be our main focus. Notice that these RGE effects are always present in any SUSY seesaw model, independently of the presence of the other sources in (i) and (ii). In this sense, the effects considered here are independent of the particular mechanism of SUSY breaking and mediation. These model dependent effects in (i) and (ii) can always be added to our results in the relevant cases.

¹ We assume R-parity conservation throughout the present work.

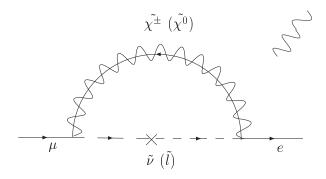


FIG. 1: The diagrams contributing to $\mu \to e, \gamma$ decays

Irrespective of the source, LFV at the weak scale can be parametrised in a model-independent manner in terms of a mass insertion (MI), Δ_{ij}^l , the flavour violating off-diagonal entry appearing in the slepton mass matrix². These MI are further subdivided into LL/LR/RL/RR types, labelled by the chirality of the corresponding SM fermions³. Depending on the model, one or several of these types of MI can simultaneously be present at the weak scale. In the presence of any of these parameters, 1-loop diagrams mediated by gauginos, higgsinos (neutral and 3 charged fermionic partners of gauge and Higgs bosons) and sleptons lead to lepton flavour violating processes such as $\mu \to e + \gamma$, $\mu \to 3e$, $\mu \to e$ conversion in nuclei, etc (an example diagram is shown in Fig.1). The strength of these processes crucially depends on the MI factor $\delta^l_{ij} \equiv \Delta^l_{ij}/m^2_{\tilde{l}}$, where $m^2_{\tilde{l}}$ is the average slepton mass. For $|\delta| < 1$, which is expected to be the case for most models, one can always use the MI approximation [15, 19] to compute the amplitudes of the relevant processes. Such computations have been done long ago, considering the neutral gaugino diagrams [6, 7]. It has been realised later that, in addition to the flavour violating LL/RR MI, considering the Higgsinos/gaugino mixing, as well as the flavour diagonal left-right mixing in the slepton mass matrix, can significantly enhance the amplitudes of these processes at large tan β [20]. These computations have since then been updated by Hisano-Nomura [21] and Masina-Savoy [22], including this mixing as well as the charged gaugino/higgsino contribution⁴. Taking the tan β factor into account, the branching ratio of $l_i \rightarrow l_i, \gamma$ for the dominant LL MI is roughly given by:

$$BR(l_j \to l_i \gamma) \approx \frac{\alpha^3 |\delta_{ij}^l|^2}{G_F^2 m_{SUSY}^4} \tan^2 \beta, \tag{1}$$

where $m_{\rm SUSY}$ represents the typical supersymmetry breaking mass such as the gaugino/slepton mass. For large $|\delta| \sim 1$ or for many δ 's present simultaneously, it is instructive to diagonalise the slepton mass matrix and evaluate the precise amplitudes in the mass-eigenstate basis. A complete computation in this basis has been presented in [23] for several LFV processes such as $l_j \to l_i + \gamma$; $l_j \to 3l_i$; $\mu \to e$ conversion in nuclei. The processes discussed so far are the ones mediated by neutralino and chargino sector. However, Higgs bosons (h^0, H^0, A^0) are also sensitive to flavour violation and mediate processes such as $\mu \to e$ conversion [25], $\tau \to 3\mu$ [26], $\tau \to \mu\eta$ [27]. The amplitudes of these processes are sensitive to a higher degree in tan β than the chargino/neutralino

 $^{^{2}}$ In the basis where the charged lepton mass matrix is diagonal.

 $^{^{3}}$ i, j, k denote generation indices throughout the present work.

⁴ Another important feature is that the interference between various contributions could lead to suppressed amplitudes in some regions of the parameter space [21, 22, 23]. This typically occurs for RR type MI as long as universality in the gaugino masses is maintained at the high scale. Although in a completely generic situation without any universal boundary conditions, such cancellations can occur for LL type MI also [24].

ones (the BRs grow as $(\tan \beta)^6$, though they are suppressed by additional Yukawa couplings) and thus could lead to large branching fractions at large tan β . Detailed studies on the $\mu \to \tau$ sector have been presented in the literature [28].

In the rest of the review, we will consider the decay $l_j \to l_i$, γ as the prototype signature of the lepton flavour violation. We will discuss several ideas put forward in the literature on the sensitivity of this decay process in determining the seesaw parameter space as well as the supersymmetric parameter space. Before this, we will briefly review the SUSY seesaw mechanism and the generation of flavour violation in this model.

III. SUPERSYMMETRIC SEESAW AND LEPTONIC FLAVOUR VIOLATION

The seesaw mechanism can be incorporated in the Minimal Supersymmetric Standard Model in a manner similar to what is done in the Standard Model, by adding right-handed neutrino superfields to the MSSM superpotential:

$$W = h_{ij}^{u} Q_{i} u_{j}^{c} H_{2} + h_{ii}^{d} Q_{i} d_{i}^{c} H_{1} + h_{ii}^{e} L_{i} e_{i}^{c} H_{1} + h_{ij}^{\nu} L_{i} \nu_{i}^{c} H_{2} + M_{R_{ii}} \nu_{i}^{c} \nu_{i}^{c} + \mu H_{1} H_{2},$$

$$(2)$$

where we are in the basis of diagonal charged lepton, down quark and right-handed Majorana mass matrices. M_R represents the (heavy) Majorana mass matrix for the right-handed neutrinos. Equation (2) leads to the standard seesaw formula for the (light) neutrino mass matrix

$$\mathcal{M}_{\nu} = -h^{\nu} M_R^{-1} h^{\nu} T v_2^2, \tag{3}$$

where v_2 is the vacuum expectation value (VEV) of the up-type Higgs field, H_2 . Under suitable conditions on h^{ν} and M_R , the correct mass splittings and mixing angles in \mathcal{M}_{ν} can be obtained. Detailed analyses deriving these conditions are already present in the literature [29].

Following the discussion in the previous section, we will assume that the mechanism that breaks supersymmetry and conveys it to the observable sector at the high scale $\sim M_{\rm P}$ is flavour blind, as in mSUGRA. However, this flavour blindness is not protected down to the weak scale [6] ⁵. The slepton mass matrices are no longer invariant under RG evolution from the super large scale where supersymmetry is mediated to the visible sector down to the seesaw scale, as the flavour violation present in the neutrino Dirac Yukawa couplings h^{ν} is now 'felt' by the slepton mass matrices in the presence of heavy right-handed neutrinos.

The weak-scale flavour violation so generated can be obtained by solving the RGEs for the slepton mass matrices from the high scale to the scale of the right-handed neutrinos. Below this scale, the running of the FV slepton mass terms is RG-invariant as the right-handed neutrinos decouple from the theory. For the purpose of illustration, a leading log estimate can easily be obtained for these equations⁶. Assuming the flavour blind mSUGRA specified by the high-scale parameters, m_0 , the common scalar mass, A_0 , the common trilinear coupling, and $M_{1/2}$, the universal gaugino mass, the flavour violating entries in these mass matrices at the weak scale are given as:

$$(\Delta_{ij}^l)_{\rm LL} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} \sum_k (h_{ik}^{\nu} h_{jk}^{\nu*}) \ln \frac{M_X}{M_{R_k}},$$
 (4)

⁵ This is always true in a gravity mediated supersymmetry breaking model, but it also applies to other mechanisms under some specific conditions [30, 31].

⁶ Within mSUGRA, the leading log approximation works very well for most of the parameter space, except for regions of large $M_{1/2}$ and low m_0 . The discrepancy with the exact result increases with low tan β [32].

where h^{ν} are given in the basis of diagonal charged lepton masses and diagonal Majorana right-handed neutrino mass matrix M_R , and M_X is the scale at which soft terms appear in the lagrangian. Given this, the branching ratios for LFV rare decays $l_j \to l_i$, γ can be roughly estimated using Eq. (1). From above it is obvious that the amount of lepton flavour violation generated by the SUSY seesaw at the weak scale crucially depends on the flavour structure of h^{ν} and M_R , shown in Eq. (2), the 'new' sources of flavour violation not present in the MSSM. If either the neutrino Yukawa couplings or the flavour mixings present in h^{ν} are very tiny, the strength of LFV will be significantly reduced. Further, if the right-handed neutrino masses were heavier than the supersymmetry breaking scale (as in GMSB models) they would decouple from the theory before the SUSY soft breaking matrices enter into play and hence these effects would vanish.

A. Model Independent Expectations for LFV?

A crucial feature of the seesaw mechanism is that it has a larger number of parameters than those relevant for neutrino masses and mixings. This would inhibit us in computing model independent expectations for, say, $BR(\mu \to e, \gamma)$, given that the supersymmetric seesaw mechanism works at the high scale. In fact, even after having a complete knowledge of the entire neutrino mass matrix elements, \mathcal{M}_{ν} as well as the heavy neutrino Majorana mass matrix eigenvalues M_{R_k} , it is still not sufficient to completely determine the rest of the seesaw parameters, namely the neutrino Dirac Yukawa coupling matrix h^{ν} . This is best illustrated in the parametrisation given by Casas and Ibarra [33]; starting from the seesaw formula, Eq. (3), one can derive h^{ν} in terms of low energy parameters as:

$$h^{\nu} = U_{\text{PMNS}}^{\star} \, \mathcal{D}_{\sqrt{\mathcal{M}_{\nu}}} \, R^T \, \mathcal{D}_{\sqrt{M_R}} \, \frac{1}{v_2}, \tag{5}$$

where $U_{\rm PMNS}$ is the Pontecorvo-Maki-Nakagawa-Sakata leptonic mixing matrix, $\mathcal{D}_{\sqrt{\mathcal{M}_{\nu}}}$ is the square root of the diagonal matrix of light neutrino mass eigenvalues, $\mathcal{D}_{\sqrt{M_R}}$ is the square root of the diagonal matrix of heavy neutrino mass eigenvalues and R is an arbitrary complex orthogonal matrix such that $RR^T=1$. The matrix R parameterises our ignorance of the neutrino Yukawas in spite of the complete knowledge of the neutrino mass matrix. Though it is difficult to give a physical definition to 'R' it is important to note that it can have physical consequences for lepton flavour violation (as well as leptogenesis), even if the neutrino masses and mixings are already accounted for. For lepton flavour violation, the relevant part of the seesaw parameter space are the entries in the matrix $h^{\nu}h^{\nu}$ which are now given by:

$$h^{\nu} h^{\nu \dagger} = U_{PMNS}^{\star} \mathcal{D}_{\sqrt{\mathcal{M}_{\nu}}} R^{T} \mathcal{D}_{M_{R}} R^{\star} \mathcal{D}_{\sqrt{\mathcal{M}_{\nu}}} U_{PMNS}^{T} \frac{1}{v_{2}^{2}}.$$
 (6)

R can be parameterised by three angles and three phases. Flavour violation is now determined in terms of the angles and phases in R, U_{PMNS} , \mathcal{M}_{ν} and M_R . For a given neutrino spectrum, U_{PMNS} and neutrino mass eigenvalues are (approximately) known. However, the branching ratios can now be either enhanced or suppressed depending on the parameters in R. Choosing M_R to be either completely hierarchical or degenerate leads to a reduction in the number of parameters affecting the branching ratios. It is further reduced if R is chosen to be real. For example, if M_R is completely hierarchical and R is real, basically only one angle would affect the branching ratios. Similarly, when M_R is degenerate and R is real, the branching ratios are independent of R. Lepton flavour violating rates for various cases of interest have been analysed in [33], special cases in [34, 35].

As mentioned earlier, flavour violation at the weak scale can be treated independently of the source in terms of the mass insertion parameters Δ_{ij}^l . In the seesaw model, these Δ parameters

are generated by RGE evolution and are thus proportional to the neutrino Yukawa couplings. In mSUGRA, the relation between these two parameters is given as⁷:

$$(\Delta_{ij}^l)_{\rm LL} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} C_{ij},$$
 (7)

where C_{ij} is defined to contain all the information from neutrino Yukawa couplings, specifically, the left-mixing angles and the eigenvalues [36, 37]; it is given by:

$$C_{ij} = [h^{\nu}h^{\nu}]_{ij} \log \frac{M_X}{M_R}. \tag{8}$$

 C_{ij} is the part of the 'seesaw' parameter space that can be probed by LFV experiments. The various upper bounds on δ_{ij}^l from the experimental limits on $l_j \to l_i$, γ decay rates and other LFV processes can now be converted to upper bounds on off-diagonal C_{ij} parameters [36] for each point in the SUSY breaking parameter space. In Table I, we present limits on $(\delta_{ij}^l)_{\rm LL}$ from the present and upcoming limits on $l_j \to l_i + \gamma$ processes [38, 39].

$(\delta^l)_{ m LL}$	Present limits	Upcoming limits
12	3×10^{-4}	8.5×10^{-6}
13	0.09	0.02
23	0.09	0.02

TABLE I: Limits on $\delta_{\rm LL}$ parameters derived in the mSUGRA model for tan β =10. The average slepton mass is chosen to be around 400 GeV and $m_0 \sim 400$ GeV. Our results agree with those presented in Refs. [21, 22].

The implied bounds on C_{ij} can in turn be used to study constraints on neutrino Dirac Yukawa entries, for example, arising from a flavour model. The sensitivity of LFV experiments in probing the seesaw parameter space has been analysed by Ellis and collaborators [37]. Choosing well defined points in supersymmetric parameter space, such as the popular Benchmark/Snowmass points, scatter plots of C_{ij} parameter space⁸ probed by several LFV processes such as $\mu \to e, \gamma$ are presented.

B. CP violation in the lepton sector

In addition to LFV, the SUSY seesaw can be responsible for several CP violating phenomena at both high and low energies in the leptonic sector. This can be facilitated by large amount of complex phases present in the seesaw couplings h^{ν} and/or M_R . We list some of these phenomena below. (i) Leptogenesis [40]

- (ii) CP violation in neutrino oscillations [41]
- (iii) CP violation in $\Delta L = 2$ processes, such as neutrinoless double beta decay [42]
- (iv) CP violation in $\Delta L = 1$ processes, such as CP violation in $\mu \to 3e$ etc [43, 44].
- (v) Leptonic EDMs [45, 46, 47]
- (vi) Slepton oscillations and CP violation in the sleptonic sector [48].

Since the source of both LFV and some of the above phenomena, say, leptogenesis, is the same

⁷ Note that only LL type Δ s are generated in the SUSY seesaw mechanism as long as one sticks to the superpotential given in Eq. (2).

⁸ C_{ij} is denoted as H_{ij} in Ref. [37].

set of parameters in the lagrangian, some amount of correlation can be expected between these phenomena for some regions of the entire seesaw parameter space. However, such a correlation is in general not guaranteed, even if the two phenomena exist simultaneously. Such a study was carried out, by random scanning of the parameter space, by Ellis and collaborators in the case of hierarchical heavy neutrinos in Ref. [49] and for degenerate heavy neutrinos in Ref. [50]. Implications of low energy observables, specifically LFV, for leptogenesis and vice versa can also be studied by using the R-parametrisation introduced above for h^{ν} , or any of the other suitable parametrisations of a complex generic matrix, h^{ν} . Such studies have been carried out in [51, 52, 53, 54]. Other related studies are [55, 56]. A detailed analysis of correlations between all the leptonic phenomena is clearly beyond the scope of the present work.

C. From LFV to seesaw parameters

As we have seen so far, measuring the neutrino mass parameters $\Delta m^2_{\rm atm},\,\Delta m^2_{\odot}$ and the mixing angles θ_{ij} more precisely would not lead us to any information on either the left-mixing or the rightmixing angles present in h^{ν} or its eigenvalues. While it also holds true within the Standard Model seesaw, the major advantage in the supersymmetric seesaw is the rich FCNC and CP violating phenomena associated with it. A natural question that follows is: Can we determine all the seesaw parameters by purely low energy experiments? At this juncture, frankly, the question looks a bit ambitious given that we still have not yet completed the determination of the neutrino mixing matrix angles and phases as well as $sg(\Delta m^2_{\rm atm})$. However, taking into consideration that (i) we have the possibility of low energy supersymmetry being observable at the LHC, (ii) being probed in great detail at linear colliders, (iii) having improved sensitivity at the upcoming facilities such as MEGA and super-B/charm factories, and (iv) improved determination of neutrino mass parameters at JPARC and long-base-line experiments, such an analysis may be required in the near future. As we will discuss later on, assuming a 'best case' scenario, the first hint of SUSY seesaw might in fact come from a $\mu \to e, \gamma$ decay, probably even before the advent of the LHC [57]. Later, more detailed evidence might pile up. The authors of Ref. [58] have presented a discussion on evaluating the seesaw parameters from low energy observables. It has been shown that although there is a one-to-one correspondence between measurables at the weak-scale and the high-scale parameters, in practice, even in the simplest supersymmetric theories, such an analysis could be formidable to achieve. This is particularly true regarding the additional phases present in the sneutrino/slepton mass matrices. However, we continue to remain optimistic and let the future decide.

IV. WHAT COULD BE THE NEUTRINO YUKAWA COUPLINGS

In the above, we have seen that due to our lack of knowledge on the neutrino Yukawa couplings, our predictions for lepton flavour violation are not effective enough in a purely bottom-up approach. This can be considered as a limitation of the model we have been working with, namely, the MSSM with right-handed neutrinos. A possible way out then could be to enlarge the SM gauge group to a much larger group, such as a Grand Unified Theory. In these models, typically quark and leptons sit in the same multiplets, leading to relations between their Yukawa couplings. One would expect that similar correlations would occur when the seesaw mechanism is incorporated within a GUT. Under the simplest of the GUT groups, 'SU(5)', however, the right-handed neutrinos remain singlets and the problem with LFV predictions still persists. One needs at least a group encompassing the Pati–Salam SU(4)_c symmetry, like an SO(10) model. Here we review one such example.

GUT models: An SO(10) example

In the SO(10) gauge theory, all the known fermions and the right-handed neutrinos are unified in a single representation of the gauge group, the 16. The product of two 16 matter representations can only couple to 10, 120 or 126 representations, which can be formed by either a single Higgs field representation or a non-renormalisable product of representations of several Higgs fields. In either case, the Yukawa matrices resulting from the couplings to 10 and 126 are complex-symmetric, whereas they are antisymmetric when the couplings are to the 120. Thus, the most general SO(10) superpotential relevant to fermion masses can be written as

$$W_{SO(10)} = h_{ij}^{10} 16_i \ 16_j \ 10 + h_{ij}^{126} 16_i \ 16_j \ 126 + h_{ij}^{120} 16_i \ 16_j \ 120, \tag{9}$$

where i, j refer to the generation indices. In terms of the SM fields, the Yukawa couplings relevant for fermion masses are given by [59] 9:

16 16 10
$$\supset 5 (uu^c + \nu \nu^c) + \bar{5} (dd^c + ee^c),$$

16 16 126 $\supset 1 \nu^c \nu^c + 15 \nu \nu + 5 (uu^c - 3 \nu \nu^c) + \bar{45} (dd^c - 3 ee^c),$
16 16 120 $\supset 5 \nu \nu^c + 45 uu^c + \bar{5} (dd^c + ee^c) + \bar{45} (dd^c - 3 ee^c),$ (10)

where we have specified the corresponding SU(5) Higgs representations for each of the couplings and all the fermions are left handed fields. The resulting mass matrices can be written as

$$M^u = M_{10}^5 + M_{126}^5 + M_{120}^{45}, (11)$$

$$M_{LR}^{\nu} = M_{10}^5 - 3 M_{126}^5 + M_{120}^5, (12)$$

$$M^{d} = M_{10}^{\bar{5}} + M_{126}^{\bar{45}} + M_{120}^{\bar{5}} + M_{120}^{\bar{45}}, \tag{13}$$

$$M^{d} = M_{10}^{\bar{5}} + M_{126}^{\bar{45}} + M_{120}^{\bar{5}} + M_{120}^{\bar{45}},$$

$$M^{e} = M_{10}^{\bar{5}} - 3M_{126}^{\bar{45}} + M_{120}^{\bar{5}} - 3M_{120}^{\bar{45}},$$

$$(13)$$

$$M_{LL}^{\nu} = M_{126}^{15}, \tag{15}$$

$$M_R^{\nu} = M_{126}^1. {16}$$

A simple analysis of the above mass matrices leads us to the following result: At least one of the Yukawa couplings in $h^{\nu} = v_u^{-1} M_{LR}^{\nu}$ has to be as large as the top Yukawa coupling [61]. This result holds true in general, independently of the choice of the Higgses responsible for the masses in Eqs. (11), (12), provided that no accidental fine-tuned cancellations of the different contributions in Eq. (12) are present. If contributions from the 10's solely dominate, h^{ν} and h^{u} would be equal. If this occurs for the 126's, then $h^{\nu} = -3 h^{\mu}$ [62]. In case both of them have dominant entries, barring a rather precisely fine-tuned cancellation between M_{10}^5 and M_{126}^5 in Eq. (12), we expect at least one large entry to be present in h^{ν} . A dominant antisymmetric contribution to top quark mass due to the 120 Higgs is phenomenologically excluded, since it would lead to at least a pair of heavy degenerate up quarks.

Apart from sharing the property that at least one eigenvalue of both M^u and M_{LR}^{ν} has to be large, for the rest it is clear from (11) and (12) that these two matrices are not aligned in general, and hence we may expect different mixing angles appearing from their diagonalisation. This freedom is removed if one sticks to particularly simple choices of the Higgses responsible for up quark and neutrino masses. A couple of remarks are in order here. Firstly, note that in general there can be an additional contribution, Eq. (15), to the light neutrino mass matrix, independent

⁹ Recently, SO(10) couplings have also been evaluated for various renormalisable and non-renormalisable couplings

of the canonical seesaw mechanism. Taking into consideration also this contribution leads to the so-called Type-II seesaw formula [63]. Secondly, the correlation between neutrino Dirac Yukawa coupling and the top Yukawa is in general independent of the type of seesaw mechanism, and thus holds true irrespective of the light-neutrino mass structure.

B. What the neutrino Yukawa mixing matrices could be

Within the SO(10) example, we have seen above that the amount of mixing present in the neutrino Dirac Yukawa couplings, h^{ν} depends on the type and number of Higgs representations there are in the theory. Motivated by the flavour structure of Standard Model fermions (including neutrino mixing angles), one can imagine two cases of mixings to be present in h^{ν} . The first one corresponds to a case where the mixing present in h^{ν} is small and CKM-like. We will call this case, 'the minimal case'. This is typical of models in which quarks and leptons have the same mixing angles at the high scale. The required large mixing angles in the light-neutrino sector is purely resultant of the seesaw mechanism. As a second case, we consider scenarios where the mixing in h^{ν} is no longer small, but large like the observed PMNS mixing. In this case, the heavy neutrino mass matrix only plays the role of a large scale due to which light neutrino masses are suppressed. We will call this case the 'the maximal case' ¹⁰. These two cases serve as 'benchmark' scenarios for seesaw induced lepton flavour violation in SUSY SO(10). Similar studies of these two extreme cases have also been considered in [23, 37].

1. Minimal case: A model for CKM mixings

The minimal Higgs spectrum to obtain phenomenologically viable mass matrices includes two 10-plets, one coupling to the up-sector and the other to the down-sector. In this way it is possible to obtain the required CKM mixing [64] in the quark sector. The SO(10) superpotential is now given by

$$W_{SO(10)} = \frac{1}{2} h_{ij}^{u,\nu} 16_i \ 16_j \ 10_u + \frac{1}{2} h_{ij}^{d,e} 16_i \ 16_j \ 10_d + \frac{1}{2} h_{ij}^R \ 16_i \ 16_j \ 126.$$
 (17)

We further assume that the **126** dimensional Higgs field gives Majorana mass *only* to the right-handed neutrinos. An additional feature of the above mass matrices is that all of them are *symmetric*. From here, it is clear that the following mass relations hold between the quark and leptonic mass matrices at the GUT scale¹¹:

$$h^u = h^\nu \quad ; \quad h^d = h^e. \tag{18}$$

In the basis where charged lepton masses are diagonal, we have

$$h^{\nu} = V_{\text{CKM}}^{T} h_{diag}^{u} V_{\text{CKM}}. \tag{19}$$

The large couplings in $h^{\nu} \sim \mathcal{O}(h_t)$ induce significant off-diagonal entries in $m_{\tilde{L}}^2$ through the RG evolution between $M_{\rm GUT}$ and the scale of the right-handed Majorana neutrinos ¹², M_{R_i} . The

¹⁰ Note that these two cases are the two examples of how one can generate large mixing angles for light-neutrino mass matrices from the seesaw mechanism [29].

¹¹ Clearly this relation cannot hold for the first two generations of down quarks and charged leptons. One expects, small corrections due to non-renormalisable operators or suppressed renormalisable operators [65] to be invoked.

 $^{^{12}}$ Typically one has different mass scales associated with different right-handed neutrino masses.

induced off-diagonal entries relevant to $l_j \to l_i, \gamma$ are of the order of:

$$(m_{\tilde{L}}^2)_{21} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 V_{td} V_{ts} \ln \frac{M_{\text{GUT}}}{M_{R_3}} + \mathcal{O}(h_c^2),$$
 (20)

$$(m_{\tilde{L}}^2)_{32} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 V_{tb} V_{ts} \ln \frac{M_{\text{GUT}}}{M_{R_3}} + \mathcal{O}(h_c^2),$$
 (21)

$$(m_{\tilde{L}}^2)_{31} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 V_{tb} V_{td} \ln \frac{M_{\text{GUT}}}{M_{R_3}} + \mathcal{O}(h_c^2).$$
 (22)

In these expressions, the CKM angles are small but one would expect the presence of the large top Yukawa coupling to compensate such a suppression. The required right-handed neutrino Majorana mass matrix, consistent with both the observed low energy neutrino masses and mixings as well as with CKM-like mixings in h^{ν} is easily determined from the seesaw formula defined at the scale of right-handed neutrinos ¹³

$$M_R = V_{\text{CKM}} h_{diag}^u V_{\text{CKM}}^T m_{\nu}^{-1} V_{\text{CKM}} h_{diag}^u V_{\text{CKM}}^T, \tag{23}$$

where we have used Eq. (19) for h^{ν} . For hierarchical neutrino mass spectrum, $m_{\nu_3} \approx \sqrt{\Delta m_{\rm atm}^2}$, $m_{\nu_2} \approx \sqrt{\Delta m_{\odot}^2}$, and $m_{\nu_1} \ll \sqrt{\Delta m_{\odot}^2}$ and for a nearly bi-maximal $U_{\rm PMNS}$, it is straightforward to see that the right handed neutrino mass eigenvalues are given by:

$$M_{R_3} \approx \frac{m_t^2}{4 m_{\nu_1}} \; ; \; M_{R_2} \approx \frac{m_c^2}{4 m_{\nu_1}} \; ; \; M_{R_1} \approx \frac{m_u^2}{2 m_{\nu_1}} \; .$$
 (24)

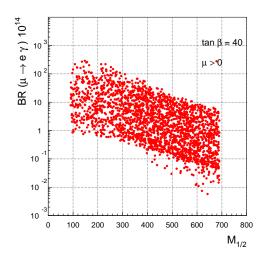


FIG. 2: The scatter plots of branching ratios of $\mu \to e, \gamma$ decays as a function of $M_{1/2}$ are shown for the (minimal) CKM case for tan $\beta = 40$. Results do not alter significantly with the change of sign(μ).

The Br $(\mu \to e, \gamma)$ is now predictable in this case. Considering mSUGRA boundary conditions, we compute these branching ratios numerically. In Fig. 2) we show the scatter plots (in mSUGRA

¹³ The neutrino masses and mixings here are defined at M_R . Radiative corrections can significantly modify the neutrino spectrum from that of the weak scale [66]. This is more true for the degenerate spectrum of neutrino masses [67] and for some specific forms of h^{ν} [68]. For our present discussion, with hierarchical neutrino masses and up-quark like neutrino Yukawa matrices, we expect these effects not to play a very significant role.

parameter space m_0 , A_0 , $M_{1/2}$) for $\mathrm{BR}(\mu \to e, \gamma)$ for the CKM case and $\tan \beta = 40$. We see that reaching a sensitivity of 10^{-14} for $\mathrm{BR}(\mu \to e\gamma)$ would allow us to probe the SUSY spectrum completely up to $M_{1/2} = 300$ GeV (notice that this corresponds to gluino and squark masses of order 750 GeV) and would still probe large regions of the parameter space up to $M_{1/2} = 700$ GeV. Thus, in summary, though the present limits on $\mathrm{BR}(\mu \to e, \gamma)$ would not induce any significant constraints on the supersymmetry-breaking parameter space, an improvement in the limit to $\sim \mathcal{O}(10^{-14})$, as foreseen, would start imposing non-trivial constraints especially for the large $\tan \beta$ region.

A further comment on the 'minimal' mixing case is in order. Strictly speaking this is not the 'minimalest' mixing possible in the Dirac neutrino Yukawa couplings. It has been shown in models where the right-handed neutrinos attain their masses through Yukawa couplings, one can essentially set the Dirac neutrino Yukawa mixing to be zero and the entire mixing comes from the right handed neutrino sector [69]. Such a situation can be realised within left-right symmetric models, with or without an SO(10) embedding. The renormalisation group flow is different in this case and the LFV is now related to Yukawa couplings of the right-handed neutrino mass matrix in an indirect manner [69, 70].

2. Maximal case: A way for PMNS mixings

The minimal SO(10) model presented in the previous subsection would inevitably lead to small mixing in h^{ν} . In fact, with two Higgs fields in symmetric representations, giving masses to the up-sector and the down-sector separately, it would be difficult to avoid the small CKM-like mixing in h^{ν} . To generate mixing angles larger than CKM angles, asymmetric mass matrices have to be considered. In general, it is sufficient to introduce asymmetric textures either in the up-sector or in the down-sector. In the present case, we assume that the down-sector couples to a combination of Higgs representations (symmetric and antisymmetric)¹⁴ Φ , leading to an asymmetric mass matrix in the basis where the up-sector is diagonal. As we will see below, this would also require that the right-handed Majorana mass matrix be diagonal in this basis. We have:

$$W_{SO(10)} = \frac{1}{2} h_{ii}^{u,\nu} 16_i 16_i 10^u + \frac{1}{2} h_{ij}^{d,e} 16_i 16_j \Phi + \frac{1}{2} h_{ii}^R 16_i 16_i 126 , \qquad (25)$$

where the **126**, as before, generates only the right-handed neutrino mass matrix. To study the consequences of these assumptions, we see that at the level of SU(5), we have

$$W_{SU(5)} = \frac{1}{2} h_{ii}^{u} 10_{i} 10_{i} 5_{u} + h_{ii}^{\nu} \bar{5}_{i} 1_{i} 5_{u} + h_{ij}^{d} 10_{i} \bar{5}_{j} \bar{5}_{d} + \frac{1}{2} M_{ii}^{R} 1_{i} 1_{i},$$
 (26)

where we have decomposed the 16 into $10 + \bar{5} + 1$ and 5_u and $\bar{5}_d$ are components of 10_u and Φ respectively. To have large mixing $\sim U_{\rm PMNS}$ in h^{ν} we see that the asymmetric matrix h^d should now be able to generate both the CKM mixing as well as PMNS mixing. This is possible if

$$V_{\text{CKM}}^T h^d U_{\text{PMNS}}^T = h_{diag}^d. (27)$$

This would mean that the 10 that contains the left-handed down-quarks would be rotated by the CKM matrix whereas the $\bar{5}$ that contains the left handed charged leptons would be rotated by

¹⁴ The couplings of Φ in the superpotential can be either renormalisable or non-renormalisable. See [71] for a non-renormalisable example.

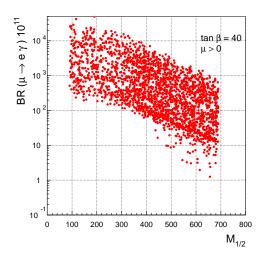


FIG. 3: The scatter plots of branching ratios of $\mu \to e, \gamma$ decays as a function of $M_{1/2}$ are shown for the (maximal) PMNS case for tan $\beta = 40$. The results do not alter significantly with the change of sign(μ).

the $U_{\rm PMNS}$ matrix to go into their respective mass bases [71, 72]. Thus we have, in analogy with the previous subsection, the following relations to hold true in the basis where charged leptons and down quarks are diagonal:

$$h^u = V_{\text{CKM}} h^u_{diag} V^T_{\text{CKM}} , \qquad (28)$$

$$h^{\nu} = U_{\text{PMNS}} h^{u}_{diag}. \tag{29}$$

Using the seesaw formula of Eqs. (3) and (29), we have

$$M_R = diag\{\frac{m_u^2}{m_{\nu_1}}, \frac{m_c^2}{m_{\nu_2}}, \frac{m_t^2}{m_{\nu_3}}\}.$$
 (30)

We now turn our attention to lepton flavour violation in this case. The branching ratio, $BR(\mu \to e, \gamma)$ would now be dependent on:

$$[h^{\nu}h^{\nu}]_{21} = h_t^2 U_{\mu 3} U_{e3} + h_c^2 U_{\mu 2} U_{e2} + \mathcal{O}(h_u^2). \tag{31}$$

It is clear from the above that in contrast to the CKM case, the dominant contribution to the off-diagonal entries depends on the unknown magnitude of the element U_{e3} [73]. If U_{e3} is very close to its present limit ~ 0.2 [74], the first term on the RHS of the Eq. (31) would dominate. Moreover, this would lead to large contributions to the off-diagonal entries in the slepton masses with $U_{\mu 3}$ of $\mathcal{O}(1)$. We have :

$$(m_{\tilde{L}}^2)_{21} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 U_{e3} U_{\mu 3} \ln \frac{M_{\text{GUT}}}{M_{R_2}} + \mathcal{O}(h_c^2).$$
 (32)

The above contribution is larger than the CKM case by a factor of $(U_{\mu3}U_{e3})/(V_{td}V_{ts}) \sim 140$ compared with the CKM case. From Eq. (1) we see that it would mean about a factor 10^4 times larger than the CKM case in BR($\mu \to e, \gamma$). In case U_{e3} is very small, *i.e* either zero or (h_c^2/h_t^2) $U_{e2} \sim 4 \times 10^{-5}$, the second term (h_c^2) in Eq. (31) would dominate. However the off-diagonal contribution in slepton masses, now being proportional to charm Yukawa could be much smaller, in fact even smaller than the CKM contribution by a factor

$$\frac{h_c^2 \ U_{\mu 2} \ U_{e2}}{h_t^2 \ V_{td} \ V_{ts}} \sim 7 \times 10^{-2}.$$
 (33)

If U_{e3} is close to its present limit, the current bound on $BR(\mu \to e, \gamma)$ would already be sufficient to produce stringent limits on the SUSY mass spectrum. Similar U_{e3} dependence can be expected in the $\tau \to e$ transitions where the off-diagonal entries are given by:

$$(m_{\tilde{L}}^2)_{31} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 U_{e3} U_{\tau 3} \ln \frac{M_{\text{GUT}}}{M_{R_3}} + \mathcal{O}(h_c^2).$$
 (34)

The $\tau \to \mu$ transitions are instead U_{e3} -independent probes of SUSY, whose importance was first pointed out in Ref. [75]. As in the rest of the cases, the off-diagonal entry in this case is given by:

$$(m_{\tilde{L}}^2)_{32} \approx -\frac{3m_0^2 + A_0^2}{8\pi^2} h_t^2 U_{\mu 3} U_{\tau 3} \ln \frac{M_{\text{GUT}}}{M_{R_3}} + \mathcal{O}(h_c^2).$$
 (35)

In the PMNS scenario, Fig. 3 shows the plot for BR($\mu \to e, \gamma$) for tan $\beta = 40$. In this plot, the value of U_{e3} chosen is very close to the present experimental upper limit [74]. As long as $U_{e3} \stackrel{>}{\sim} 4 \times 10^{-5}$, the plots scale as U_{e3}^2 , while for $U_{e3} \stackrel{>}{\sim} 4 \times 10^{-5}$ the term proportional to m_c^2 in Eq. (32) starts dominating; the result is then insensitive to the choice of U_{e3} . For instance, a value of $U_{e3} = 0.01$ would reduce the BR by a factor of 225 and still a significant amount of the parameter space for tan $\beta = 40$ would be excluded. We further find that with the present limit on BR($\mu \to e, \gamma$), all the parameter space would be completely excluded up to $M_{1/2} = 300$ GeV for $U_{e3} = 0.15$, for any vale of tan β (not shown in the figure).

In the $\tau \to \mu \gamma$ decay the situation is similarly constrained. For $\tan \beta = 2$, the present bound of 3×10^{-7} starts probing the parameter space up to $M_{1/2} \le 150$ GeV. The main difference is that this does not depend on the value of U_{e3} , and therefore it is already a very important constraint on the parameter space of the model. In fact, for large $\tan \beta = 40$, as shown in Fig. 4, reaching the expected limit of 1×10^{-8} would be able to rule out completely this scenario up to gaugino masses of 400 GeV, and only a small portion of the parameter space with heavier gauginos would survive. In the limit $U_{e3} = 0$, this decay mode would provide a constraint on the model stronger than $\mu \to e, \gamma$, which would now be suppressed as it would contain only contributions proportional to h_c^2 , as shown in Eq. (32).

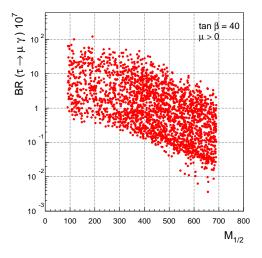


FIG. 4: The scatter plots of branching ratios of $\tau \to \mu, \gamma$ decays as a function of $M_{1/2}$ are shown for the (maximal) PMNS case for tan $\beta = 40$. The results do not alter significantly with the change of sign(μ).

In summary, in the PMNS/maximal mixing case, even the present limits from BR($\mu \to e, \gamma$) can rule out large portions of the supersymmetry-breaking parameter space, if U_{e3} is either close to

its present limit or within an order of magnitude of it (as the planned experiments might find out soon [76]). These are more severe for the large $\tan \beta$ case. In the extreme situation of U_{e3} being zero or very small $\sim \mathcal{O}(10^{-4}-10^{-5})$, $\mathrm{BR}(\tau \to \mu, \gamma)$ will start playing an important role, with its present constraints already disallowing large regions of the parameter space at large $\tan \beta$. While the above example concentrated on the hierarchical light neutrinos, similar 'benchmark' mixing scenarios have been explored in great detail, for degenerate spectra of light neutrinos, by Ref. [77], taking also in to consideration running between the Planck scale and the GUT scale.

C. Textures and other examples

While the bottom-up approach studies the possibility of 'measuring' the seesaw parameters through low energy experiments, the top-down approach gives the opportunity to study several theoretically well motivated models encompassing the seesaw mechanism. For example, a flavour symmetry based on either abelian or non-abelian family symmetries could be at work at the scales giving rise to specific patterns in all the Yukawa coupling matrices in the lagrangian including that of neutrino Dirac Yukawa couplings. Low energy flavour violation is then dependent on these patterns of the Yukawa matrices which are predictable. Several analyses of this kind have been presented in the literature mostly within a supersymmetric GUT [16, 36, 78, 79, 80]. A recent review of several textures presented in the literature can be found in [81].

Textures that tend to lead to large left-mixing in h^{ν} are typically prone to constraints from the present limits on $\mu \to e + \gamma$. A class of textures which goes by the name of 'lop-sided' generically predict large branching ratios [79] within the reach of experiments at MEGA. Finally, in addition to the SO(10) example presented here, there have been several other examples both within the context of SO(10) and otherwise that have been explored in the literature [82, 83].

V. SEESAW INDUCED LFV AND ASSOCIATED PHENOMENOLOGY

So far we have looked at LFV generated by a seesaw mechanism both through a bottom-up as well as top-down perspectives. In the following we will discuss the impact the generated LFV can have on the associated phenomenology of the SUSY model. We will only concentrate on a few issues, leaving out the leptonic CP violation which we have already commented on previously. Most of these correlated effects are only valid within a class of SUSY-GUT models, as such correlations cannot be constructed from a purely bottom-up approach.

A. LFV and experimental sensitivity to U_{e3}

In the maximal mixing situation, which we have discussed in the previous subsection, we have seen that the BR($\mu \to e, \gamma$) would depend crucially on the neutrino mixing matrix element U_{e3} . To illustrate the correlations between $\mu \to e + \gamma$ and the neutrino mixing angle U_{e3} , we chose specific points in the supersymmetric parameter space as given by the Snowmass collaboration [85]. These are presented in Table II.

In Fig. 2, we plot $BR(\mu \to e + \gamma)$ with respect to U_{e3}^2 or $\sin^2 \theta_{13}$. We also present the sensitivity of various future experiments probing U_{e3} and well as the expected improvements on the limits on $BR(\mu \to e + \gamma)$. It is seen that $\mu \to e\gamma$ can have a stronger sensitivity on U_{e3} if both SUSY seesaw and maximal mixing case are realised in nature [86]. A detailed study of the impact of U_{e3} on observability of SUSY at the LHC can be found in [57].

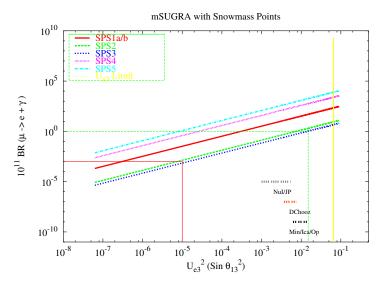


FIG. 5: The variation of the BR($\mu \to e + \gamma$) with respect to U_{e3}^2 in the maximal mixing, MNS case. Each (diagonal) line corresponds to an SPS point in mSUGRA as denoted in the legend in the figure. The sensitivity of future LFV experiments $\mathcal{O}(10^{-11}-10^{-14})$ is projected on the U_{e3}^2 axis. The ranges probed by future long-base-line experiments are shown as horizontal lines. NuI/JP represents he projected sensitivity of Nu-MI/J-PARC, DChooz, reactor-based experiments like double CHOOZ, Min/Ica/Op represents the experiments MINOS, ICARUS and OPERA [84].

	m_0	$M_{1/2}$	A_0	$\tan \beta$	$sg(\mu)$
SPS1a	100	250	-100	10	> 0
SPS1b	200	400	0	30	> 0
SPS2	1450	300	0	10	> 0
SPS3	90	400	0	10	> 0
SPS4	400	300	0	50	> 0
SPS5	150	300	-1000	5	> 0

TABLE II: SPS points for mSUGRA

B. Correlations with other SUSY search strategies

In addition to the improvements in LFV experiments, this is also going to be the decade in which we should be able to establish whether low energy supersymmetry exists or not through direct searches at the LHC [87]. On the other hand, improved astrophysical observations from experiments like WMAP [88] and Planck are going to determine the relic density of supersymmetric LSP at unprecedented accuracy. Within mSUGRA, correlations between these two search strategies have been studied [89]. Incorporating the seesaw mechanism in the model à la SO(10), would generate another discovery strategy through the lepton flavour violation channel. This is especially true when the LFV entries in the slepton mass matrices are maximised, as in the PMNS case.

We see that three main regions in the mSUGRA parameter space would survive after imposing all the present phenomenological and astrophysical (dark matter) constraints [57]¹⁵. These are: (a) The stau coannihilation regions, where the lightest stau is quasi-degenerate with the neutralino LSP and efficient stau-stau as well as stau-neutralino (co)annihilations suppress the relic density.

¹⁵ For a bottom-up analyses, see Ref. [90].

- (b) The A-pole funnel region, where the neutralino(bino)-neutralino annihilation process is greatly enhanced through a resonant s-channel exchange of the heavy neutral Higgs A and H (c) Focus point or hyperbolic branch regions, where a non-negligible higgsino fraction in the lightest neutralino is produced. In each of these regions the LFV rates emanating from the seesaw mechanism can be computed and contrasted with the sensitivity of direct searches at the LHC. Assuming the maximal mixing PMNS case, we find [57]:
 - Coannihilation regions: In these regions, which are mostly accessible at the LHC, an improvement of two orders of magnitude in the branching ratio sensitivity from the present limit, would make $\mu \to e\gamma$ visible for most of the parameter space, as long as $U_{e3} \gtrsim 0.02$, even for the low tan β region. For large tan β , independent of U_{e3} , $\tau \to \mu \gamma$ will start probing this region provided a sensitivity of $\mathcal{O}(10^{-8})$ is reached.
 - A-pole funnel regions: In these regions the LHC reach is not complete and LFV may be competitive. If $U_{e3} \gtrsim 10^{-2}$, the future $\mu \to e\gamma$ experiments, with limit of $\mathcal{O}(10^{-14})$ will probe most of the parameter space. As before, $\tau \to \mu \gamma$ will probe this region once the BR sensitivity reaches $\mathcal{O}(10^{-8})$.
 - Focus point regions: Since the LHC reach in this region is rather limited by the large m_0 and $M_{1/2}$ values, LFV could constitute a privileged road toward SUSY discovery. This would require improvements of at least a couple of orders of magnitude (or more, depending on the value of U_{e3}) of improvement on the present limit of BR($\mu \to e, \gamma$). Dark Matter (DM) searches will also have partial access to this region in future, leading to a new complementarity between LFV and the quest for the cold dark matter constituent of the Universe.

C. Seesaw induced Hadronic FCNC and CPV

So far we have seen that the SUSY version of the seesaw mechanism can lead to potentially large leptonic flavour violations, so large that they could compete even with direct searches at the LHC. If one combines these ideas of supersymmetric seesaw with those of quark–lepton unification, as in a supersymmetric GUT, one would expect that the seesaw resultant flavour effects would now be also felt in the hadronic sector, and vice versa [14, 15]. In fact, this is what happens in a SUSY SU(5) with seesaw mechanism [72], where the seesaw induced RGE effects generate flavour violating terms in the right handed squark multiplets. However, as is the case with the MSSM + seesaw mechanism, within the SU(5) model also, information from the neutrino masses is not sufficient to fix all the seesaw parameters; a large neutrino Yukawa coupling has to be assumed to have the relevant phenomenological consequences in hadronic physics, such as CP violation in $B \to \Phi K_s$ etc.

As we have already seen within the SO(10) model, a large neutrino Yukawa, of the order of that of the top quark, is almost inevitable. Using this, it has been pointed in Ref. [71], that the observed large atmospheric ν_{μ} - ν_{τ} transitions imply a potentially large $b \to s$ transitions in SUSY SO(10). In the presence of CP violating phases, this can lead to enhanced CP asymmetries in B_s and B_d decays. In particular, the still controversial discrepancy between the SM prediction and the observed $A_{\rm CP}(B_d \to \Phi K_s)$ [91] can be attributed to these effects. Interestingly, despite the severe constraints on the $b \to s$ transitions from $B \to X_s$, γ [92, 93], subsequent detailed analyses [94, 95] proved that there is still enough room for sizeble deviations from the SM expectations for CP violation in the B systems. The reader interested in various correlations in $b \to s$ transitions with all possible FV off-diagonal squark mass entries can find an exhaustive answer in Ref. [95].

Finally, let us make a short comment about possible correlations between the hadronic and leptonic FV effects in a SUSY GUT. If the FV soft breaking terms appear at a scale larger than that of the grand unification, they must be related by the GUT symmetry. This puts constraints on the boundary conditions for the running of the FV soft parameters. From this consideration, one might intuitively expect that some correlation between various leptonic and hadronic FCNC processes [38] can occur at the weak scale. If in the evolution of the sparticle masses from the grand unification scale down to the electroweak scale, one encounters seesaw physics, then the quark–lepton correlations involving the left-handed sleptons, though modified, lead to even stronger constraints on hadronic physics [38, 39]. Some other related works are [96].

VI. ALTERNATIVES TO CANONICAL SEESAW AND LFV

In the final section before conclusions, we briefly mention lepton flavour violating studies conducted in mechanisms other than the canonical seesaw. As earlier, our list is not exhaustive or detailed.

- (i) Type II seesaw: In the SO(10) example, as we have shown, one can consider a situation where the non-seesaw contribution (15) dominates over the seesaw one. Flavour violation in this case has been studied in specific models [97].
- (ii) Triplet Seesaw: The seesaw mechanism itself can be implemented with fermionic triplets instead of singlet fermions. In the supersymmetric context, the flavour violation is then proportional to the triplet couplings and the neutrino mixing angles [98].
- (iii) 3×2 Seesaw: The seesaw parameter space drastically reduces when one of the neutrinos is assumed to be too heavy and decouples from the model. The predictive power of the model is now enhanced for various low energy and high energy observables [99, 100, 101].
- (iv) X-dimensional models: In models with warped X-dimensional scenarios, large LFV is expected with either Dirac-type neutrinos [102] or Majorana-type neutrinos [103].

Flavour violation has also been computed in various other models as R-parity violation [104], seesaw models with additional leptons (or inverse seesaw) [105], left-right symmetric models [106], Standard Model seesaw with a large number of Higgs bosons [107].

VII. CONCLUSIONS

Since the discovery of neutrino masses in atmospheric neutrino oscillations, there has been a lot of activity trying to understand the seesaw mechanism and its low energy implications. On the one side, we use a bottom-up approach quantifying our ignorance on the seesaw parameter space to compute the LFV, while on the other side, in top-down approach, various models and textures are constrained by LFV. Associated implications to *B*-physics GUT models, DM abundances in SUGRA theories and implications for LHC searches have been and are being done. Though we have not discussed it here, LFV in *Z*-decays [108] and other collider processes [109] is also being investigated.

Undoubtedly, the seesaw mechanism represents (one of) the best proposal to generate small neutrino masses naturally. But, how can we make sure that this is indeed nature's choice? Even establishing the Majorana nature of the neutrinos through a positive evidence of neutrinoless double beta decay, it will be difficult to assess that such Majorana masses come from a seesaw. Indeed, as we said at the beginning, in the SM seesaw we expect very tiny charged LFV effects, probably without any chance of ever observing them. When moving to SUSY seesaw we add an important handle to our effort to establish the presence of a seesaw. In fact, as we tried to show in this work, SUSY extensions of the SM with a seesaw have a general 'tendency' to enhance (or even

strongly enhance) rare LFV processes. Hence the combination of the observation of neutrinoless double beta decay and of some charged LFV phenomenon would constitute an important clue for the assessment of SUSY seesaw in nature.

There is no doubt that after the discovery of the neutrino masses, among the indirect tests of SUSY through FCNC and CP violating phenomena, LFV processes have acquired a position of utmost relevance. It would be spectacular if, by the time the LHC observes the first SUSY particle, we could see also a muon decaying to an electron and a photon! After thirty years, we could have the simultaneous confirmation of two of the most challenging physics ideas: seesaw and low energy SUSY.

Acknowledgements: We thank all our collaborators, Marco Ciuchini, Stefano Profumo, Luca Silvestrini and Carlos Yaguna, without whom the present work would not have been possible. Diagram by Jaxodraw [110].

- [1] M. L. Brooks et al. [MEGA Collaboration], Phys. Rev. Lett. 83 (1999) 1521 [arXiv:hep-ex/9905013].
- [2] K. Abe et al. [Belle Collaboration], arXiv:hep-ex/0310029.
- [3] K. Inami, for the Belle Colloboration, Talk presented at the 19th International Workshop on Weak Interactions and Neutrinos (WIN-03) October 6th to 11th, 2003, Lake Geneva, Wisconsin, USA.
- [4] Web page: http://meg.psi.ch
- [5] P. Minkowski, Phys. Lett. B 67, 421 (1977); T. Yanagida in Proc. Workshop on Unified Theories &c., eds. O. Sawada and A. Sugamoto (Tsukuba, 1979); M. Gell-Mann, P. Ramond, and R. Slansky, in Supergravity, eds. D. Freedman et al., (North Holland 1980 Amsterdam) p. 315; S.L. Glashow, in Quarks and Leptons, Cargèse 1979, eds. M. Lévy, et al., (Plenum 1980 New York), p. 707; R. N. Mohapatra and G. Senjanovic, Phys. Rev. Lett. 44, 912 (1980).
- [6] F. Borzumati and A. Masiero, Phys. Rev. Lett. **57**, 961 (1986).
- [7] F. Gabbiani, E. Gabrielli, A. Masiero and L. Silvestrini, Nucl. Phys. B 477, 321 (1996) [arXiv:hep-ph/9604387].
- [8] For a recent review, see A. Masiero and O. Vives, Ann. Rev. Nucl. Part. Sci. 51 (2001) 161 [arXiv:hep-ph/0104027]; A. Masiero and O. Vives, New Jour. Phys. 4 (2002) 4.
- [9] For a recent review, please see D. J. H. Chung, L. L. Everett, G. L. Kane, S. F. King, J. Lykken and L. T. Wang, arXiv:hep-ph/0312378.
- [10] See for example, A. Brignole, L. E. Ibanez and C. Munoz, arXiv:hep-ph/9707209 and references therein.
- [11] E. Dudas, S. Pokorski and C. A. Savoy, Phys. Lett. B 369, 255 (1996) [arXiv:hep-ph/9509410].
- [12] S. A. Abel and G. Servant, Nucl. Phys. B 611, 43 (2001) [arXiv:hep-ph/0105262]; G. G. Ross and O. Vives, Phys. Rev. D 67, 095013 (2003) [arXiv:hep-ph/0211279].
- [13] S. F. King, I. N. R. Peddie, G. G. Ross, L. Velasco-Sevilla and O. Vives, arXiv:hep-ph/0407012.
- [14] R. Barbieri and L. J. Hall, Phys. Lett. B 338, 212 (1994) [arXiv:hep-ph/9408406]; R. Barbieri,
 L. J. Hall and A. Strumia, Nucl. Phys. B 445, 219 (1995) [arXiv:hep-ph/9501334].
- [15] L. J. Hall, V. A. Kostelecky and S. Raby, Nucl. Phys. B 267, 415 (1986);
- [16] S. F. King and I. N. R. Peddie, Nucl. Phys. B 678, 339 (2004) [arXiv:hep-ph/0307091].
- [17] See for example: L. J. Hall and Y. Nomura, Phys. Rev. D 66, 075004 (2002) [arXiv:hep-ph/0205067].
- [18] H. Abe, K. Choi, K. S. Jeong and K. i. Okumura, arXiv:hep-ph/0407005.
- [19] F. Gabbiani and A. Masiero, Nucl. Phys. B **322**, 235 (1989).
- [20] J. Hisano, T. Moroi, K. Tobe, M. Yamaguchi and T. Yanagida, Phys. Lett. B **357**, 579 (1995) [arXiv:hep-ph/9501407].
- [21] J. Hisano and D. Nomura, Phys. Rev. D 59, 116005 (1999) [arXiv:hep-ph/9810479].
- [22] I. Masina and C. A. Savoy, Nucl. Phys. B 661, 365 (2003) [arXiv:hep-ph/0211283].
- J. Hisano, T. Moroi, K. Tobe and M. Yamaguchi, Phys. Rev. D 53 (1996) 2442 [arXiv:hep-ph/9510309];
 J. Hisano, T. Moroi, K. Tobe and M. Yamaguchi, Phys. Lett. B 391, 341 (1997) [Erratum-ibid. B 397, 357 (1997)] [arXiv:hep-ph/9605296].
- [24] See for example, S. Profumo and C. E. Yaguna, Nucl. Phys. B 681, 247 (2004) [arXiv:hep-ph/0307225].

- [25] See for example, R. Kitano, M. Koike and Y. Okada, Phys. Rev. D 66, 096002 (2002) [arXiv:hep-ph/0203110]; R. Kitano, M. Koike, S. Komine and Y. Okada, Phys. Lett. B 575, 300 (2003) [arXiv:hep-ph/0308021] and references therein.
- [26] K. S. Babu and C. Kolda, Phys. Rev. Lett. 89, 241802 (2002) [arXiv:hep-ph/0206310]; A. Dedes,
 J. R. Ellis and M. Raidal, Phys. Lett. B 549, 159 (2002) [arXiv:hep-ph/0209207].
- [27] M. Sher, Phys. Rev. D 66, 057301 (2002) [arXiv:hep-ph/0207136].
- [28] A. Brignole and A. Rossi, Phys. Lett. B 566, 217 (2003) [arXiv:hep-ph/0304081]; A. Brignole and A. Rossi, arXiv:hep-ph/0404211.
- [29] Some recent reviews are: G. Altarelli and F. Feruglio, arXiv:hep-ph/0206077; arXiv:hep-ph/0306265; arXiv:hep-ph/0405048; I. Masina, Int. J. Mod. Phys. A 16, 5101 (2001) [arXiv:hep-ph/0107220]; R. N. Mohapatra, arXiv:hep-ph/0211252; arXiv:hep-ph/0306016; S. F. King, arXiv:hep-ph/0310204; A. Y. Smirnov, arXiv:hep-ph/0311259.
- [30] K. Tobe, J. D. Wells and T. Yanagida, Phys. Rev. D 69, 035010 (2004) [arXiv:hep-ph/0310148].
- [31] M. Ibe, R. Kitano, H. Murayama and T. Yanagida, arXiv:hep-ph/0403198.
- [32] S. T. Petcov, S. Profumo, Y. Takanishi and C. E. Yaguna, Nucl. Phys. B **676**, 453 (2004) [arXiv:hep-ph/0306195].
- [33] J. A. Casas and A. Ibarra, Nucl. Phys. B 618 (2001) 171 [arXiv:hep-ph/0103065].
- [34] A. Kageyama, S. Kaneko, N. Shimoyama and M. Tanimoto, Phys. Lett. B **527**, 206 (2002) [arXiv:hep-ph/0110283].
- [35] F. Deppisch, H. Paes, A. Redelbach, R. Ruckl and Y. Shimizu, Eur. Phys. J. C 28, 365 (2003) [arXiv:hep-ph/0206122].
- [36] S. Lavignac, I. Masina and C. A. Savoy, Phys. Lett. B 520, 269 (2001) [arXiv:hep-ph/0106245];
 S. Lavignac, I. Masina and C. A. Savoy, Nucl. Phys. B 633, 139 (2002) [arXiv:hep-ph/0202086].
- [37] J. R. Ellis, J. Hisano, M. Raidal and Y. Shimizu, Phys. Rev. D **66**, 115013 (2002) [arXiv:hep-ph/0206110].
- [38] M. Ciuchini, A. Masiero, L. Silvestrini, S. K. Vempati and O. Vives, Phys. Rev. Lett. 92, 071801 (2004) [arXiv:hep-ph/0307191].
- [39] M. Ciuchini, A. Masiero, P. Paradisi, L. Silvestrini, S. K. Vempati and O. Vives, to appear.
- [40] For a recent review, please see: W. Buchmuller, P. Di Bari and M. Plumacher, arXiv:hep-ph/0406014.
- [41] See for example, discussions in C. Albright *et al.*, arXiv:hep-ex/0008064; M. Blom and H. Minakata, arXiv:hep-ph/0404142.
- [42] See for example, discussions in A. de Gouvea, B. Kayser and R. N. Mohapatra, Phys. Rev. D 67, 053004 (2003) [arXiv:hep-ph/0211394]; S. Pascoli and S. T. Petcov, Phys. Lett. B 580, 280 (2004) [arXiv:hep-ph/0310003].
- [43] Y. Okada, K. i. Okumura and Y. Shimizu, Phys. Rev. D 61, 094001 (2000) [arXiv:hep-ph/9906446];
 Y. Kuno and Y. Okada, Rev. Mod. Phys. 73, 151 (2001) [arXiv:hep-ph/9909265].
- [44] J. R. Ellis, J. Hisano, S. Lola and M. Raidal, Nucl. Phys. B 621, 208 (2002) [arXiv:hep-ph/0109125].
- [45] J. R. Ellis, J. Hisano, M. Raidal and Y. Shimizu, Phys. Lett. B **528**, 86 (2002) [arXiv:hep-ph/0111324].
- [46] I. Masina, Nucl. Phys. B **671**, 432 (2003) [arXiv:hep-ph/0304299].
- [47] Y. Farzan and M. E. Peskin, arXiv:hep-ph/0405214.
- [48] N. Arkani-Hamed, H. C. Cheng, J. L. Feng and L. J. Hall, Phys. Rev. Lett. 77, 1937 (1996) [arXiv:hep-ph/9603431]; N. Arkani-Hamed, J. L. Feng, L. J. Hall and H. C. Cheng, Nucl. Phys. B 505, 3 (1997) [arXiv:hep-ph/9704205].
- [49] J. R. Ellis and M. Raidal, Nucl. Phys. B 643, 229 (2002) [arXiv:hep-ph/0206174].
- [50] J. R. Ellis, M. Raidal and T. Yanagida, Phys. Lett. B **546**, 228 (2002) [arXiv:hep-ph/0206300].
- [51] S. Davidson and A. Ibarra, Nucl. Phys. B 648, 345 (2003) [arXiv:hep-ph/0206304]; S. Davidson, JHEP 0303, 037 (2003) [arXiv:hep-ph/0302075].
- [52] G. C. Branco, R. Gonzalez Felipe, F. R. Joaquim, I. Masina, M. N. Rebelo and C. A. Savoy, Phys. Rev. D 67, 073025 (2003) [arXiv:hep-ph/0211001].
- [53] S. Pascoli, S. T. Petcov and W. Rodejohann, Phys. Rev. D 68, 093007 (2003) [arXiv:hep-ph/0302054];
- [54] S. Pascoli, S. T. Petcov and C. E. Yaguna, Phys. Lett. B **564**, 241 (2003) [arXiv:hep-ph/0301095].
- [55] J. R. Ellis, M. Raidal and T. Yanagida, arXiv:hep-ph/0303242.
- [56] K. Siyeon, Eur. Phys. J. C **30**, 55 (2003) [arXiv:hep-ph/0303077].
- [57] A. Masiero, S. Profumo, S. K. Vempati and C. E. Yaguna, JHEP **0403**, 046 (2004) [arXiv:hep-ph/0401138].

- [58] S. Davidson and A. Ibarra, JHEP **0109**, 013 (2001) [arXiv:hep-ph/0104076].
- [59] R. Barbieri, D. V. Nanopoulos, G. Morchio and F. Strocchi, Phys. Lett. B 90, 91 (1980); P. Langacker, Phys. Rept. 72 (1981) 185.
- [60] P. Nath and R. M. Syed, Phys. Lett. B 506 (2001) 68; [Erratum-ibid. B 508 (2001) 216] [arXiv:hep-ph/0103165]; P. Nath and R. M. Syed, Nucl. Phys. B 618 (2001) 138 [arXiv:hep-th/0109116]; P. Nath and R. M. Syed, Nucl. Phys. B 676 (2004) 64 [arXiv:hep-th/0310178]; B. Bajc, A. Melfo, G. Senjanovic and F. Vissani, arXiv:hep-ph/0402122; C. S. Aulakh and A. Girdhar, arXiv:hep-ph/0405074; T. Fukuyama, A. Ilakovac, T. Kikuchi, S. Meljanac and N. Okada, arXiv:hep-ph/0405300.
- [61] A. Masiero, S. K. Vempati and O. Vives, Nucl. Phys. B 649, 189 (2003) [arXiv:hep-ph/0209303]; arXiv:hep-ph/0405017.
- [62] See for example, R. N. Mohapatra and B. Sakita, Phys. Rev. D 21, 1062 (1980).
- [63] G. Lazarides, Q. Shafi and C. Wetterich, Nucl. Phys. B 181, 287 (1981); R. N. Mohapatra and G. Senjanovic, Phys. Rev. D 23, 165 (1981).
- [64] S. Dimopoulos and L. J. Hall, Phys. Lett. B 344, 185 (1995) [arXiv:hep-ph/9411273]; W. Buchmuller and D. Wyler, Phys. Lett. B 521 (2001) 291 [arXiv:hep-ph/0108216].
- [65] H. Georgi and C. Jarlskog, Phys. Lett. B 86 (1979) 297.
- [66] For a review, see P. H. Chankowski and S. Pokorski, Int. J. Mod. Phys. A 17, 575 (2002) [arXiv:hep-ph/0110249].
- J. R. Ellis and S. Lola, Phys. Lett. B 458, 310 (1999) [arXiv:hep-ph/9904279]; J. A. Casas, J. R. Espinosa, A. Ibarra and I. Navarro, Nucl. Phys. B 556, 3 (1999) [arXiv:hep-ph/9904395]; N. Haba and N. Okamura, Eur. Phys. J. C 14, 347 (2000) [arXiv:hep-ph/9906481].
- [68] S. Antusch, J. Kersten, M. Lindner and M. Ratz, Phys. Lett. B 544, 1 (2002) [arXiv:hep-ph/0206078]
 ; S. Antusch and M. Ratz, JHEP 0211, 010 (2002) [arXiv:hep-ph/0208136].
- [69] K. S. Babu, B. Dutta and R. N. Mohapatra, Phys. Rev. D 67, 076006 (2003) [arXiv:hep-ph/0211068];
- [70] B. Dutta and R. N. Mohapatra, Phys. Rev. D 68, 113008 (2003) [arXiv:hep-ph/0307163].
- [71] D. Chang, A. Masiero and H. Murayama, Phys. Rev. D 67, 075013 (2003) [arxiv:hep-ph/0205111].
- [72] T. Moroi, JHEP 0003, 019 (2000) [arXiv:hep-ph/0002208]; T. Moroi, Phys. Lett. B 493, 366 (2000) [arXiv:hep-ph/0007328]; N. Akama, Y. Kiyo, S. Komine and T. Moroi, Phys. Rev. D 64, 095012 (2001) [arXiv:hep-ph/0104263].
- [73] J. Sato, K. Tobe and T. Yanagida, Phys. Lett. B 498, 189 (2001) [arXiv:hep-ph/0010348].
- [74] M. Apollonio et al. [CHOOZ Collaboration], Phys. Lett. B 466, 415 (1999) [arXiv:hep-ex/9907037].
- [75] T. Blazek and S. F. King, Phys. Lett. B **518**, 109 (2001) [arXiv:hep-ph/0105005].
- [76] M. Goodman, arXiv:hep-ex/0404031.
- [77] J. I. Illana and M. Masip, Eur. Phys. J. C 35, 365 (2004) [arXiv:hep-ph/0307393].
- [78] An incomplete list of references: S. F. King and M. Oliveira, Phys. Rev. D 60, 035003 (1999) [arXiv:hep-ph/9804283]; J. L. Feng, Y. Nir and Y. Shadmi, Phys. Rev. D 61, 113005 (2000) [arXiv:hep-ph/9911370]. J. R. Ellis, M. E. Gomez, G. K. Leontaris, S. Lola and D. V. Nanopoulos, Eur. Phys. J. C 14, 319 (2000) [arXiv:hep-ph/9911459]; D. F. Carvalho, J. R. Ellis, M. E. Gomez and S. Lola, Phys. Lett. B 515, 323 (2001) [arXiv:hep-ph/0103256]; A. Romanino and A. Strumia, Nucl. Phys. B 622, 73 (2002) [arXiv:hep-ph/0108275]; A. Kageyama, S. Kaneko, N. Shimoyama and M. Tanimoto, Phys. Rev. D 65, 096010 (2002) [arXiv:hep-ph/0112359]; D. Falcone, Mod. Phys. Lett. A 17, 2467 (2002) [arXiv:hep-ph/0207308]; K. Hamaguchi, M. Kakizaki and M. Yamaguchi, Phys. Rev. D 68, 056007 (2003) [arXiv:hep-ph/0212172]; K. S. Babu, T. Enkhbat and I. Gogoladze, Nucl. Phys. B 678, 233 (2004) [arXiv:hep-ph/0308093]; G. G. Ross and L. Velasco-Sevilla, Nucl. Phys. B 653, 3 (2003) [arXiv:hep-ph/0208218]. S. Kaneko, M. Katsumata and M. Tanimoto, JHEP 0307, 025 (2003) [arXiv:hep-ph/0305014]. G. G. Ross, L. Velasco-Sevilla and O. Vives, Nucl. Phys. B692 (2004) 50, [arXiv:hep-ph/0401064],
 - M. Bando, S. Kaneko, M. Obara and M. Tanimoto, arXiv:hep-ph/0405071.
- [79] J. Sato and K. Tobe, Phys. Rev. D 63, 116010 (2001) [arXiv:hep-ph/0012333]. X. J. Bi, Eur. Phys. J. C 27, 399 (2003) [arXiv:hep-ph/0211236]; S. M. Barr, Phys. Lett. B 578, 394 (2004) [arXiv:hep-ph/0307372]; E. Jankowski and D. W. Maybury, Phys. Rev. D 70, 035004 (2004) [arXiv:hep-ph/0401132].
- [80] M. C. Chen and K. T. Mahanthappa, arXiv:hep-ph/0409096.
- [81] M. C. Chen and K. T. Mahanthappa, Int. J. Mod. Phys. A 18, 5819 (2003) [arXiv:hep-ph/0305088].

- [82] J. Hisano, D. Nomura and T. Yanagida, Phys. Lett. B 437, 351 (1998) [arXiv:hep-ph/9711348];
 W. Buchmuller, D. Delepine and F. Vissani, Phys. Lett. B 459, 171 (1999) [arXiv:hep-ph/9904219];
 S. Baek, T. Goto, Y. Okada and K. i. Okumura, Phys. Rev. D 63, 051701 (2001) [arXiv:hep-ph/0002141];
 J. Hisano and K. Tobe, Phys. Lett. B 510, 197 (2001) [arXiv:hep-ph/0102315];
 S. Baek, T. Goto, Y. Okada and K. i. Okumura, Phys. Rev. D 64, 095001 (2001) [arXiv:hep-ph/0104146];
 J. Hisano, arXiv:hep-ph/0209005;
 T. Blazek and S. F. King, Nucl. Phys. B 662, 359 (2003) [arXiv:hep-ph/0211368];
 C. S. Huang, T. j. Li and W. Liao, Nucl. Phys. B 673, 331 (2003) [arXiv:hep-ph/0304130];
 T. Fukuyama, T. Kikuchi and N. Okada, Phys. Rev. D 68, 033012 (2003) [arXiv:hep-ph/0304190];
 T. F. Feng, T. Huang, X. Q. Li, X. M. Zhang and S. M. Zhao, Phys. Rev. D 68, 016004 (2003) [arXiv:hep-ph/0305290];
 B. Dutta, Y. Mimura and R. N. Mohapatra, arXiv:hep-ph/0402113.
- [83] For a study of an SO(10) texture leading to interesting correlations between various flavour and CP phenomena, please see, J. C. Pati, talk presented at the Fujihara seminar, KEK Laboratory, Tsukuba, Japan (Feb 23-25,2004). We thank the author for sending us the manuscript before publication; J. C. Pati, arXiv:hep-ph/0407220.
- [84] P. Huber, M. Lindner, M. Rolinec, T. Schwetz and W. Winter, arXiv:hep-ph/0403068.
- [85] B. C. Allanach et al., in Proc. of the APS/DPF/DPB Summer Study on the Future of Particle Physics (Snowmass 2001) ed. N. Graf, Eur. Phys. J. C 25, 113 (2002) [eConf C010630, P125 (2001)] [arXiv:hep-ph/0202233].
- [86] A. Masiero, S. K. Vempati and O. Vives, contribution to Neutrino 2004.
- [87] H. Baer, C. h. Chen, F. Paige and X. Tata, Phys. Rev. D 52, 2746 (1995) [arXiv:hep-ph/9503271];
 H. Baer, C. h. Chen, F. Paige and X. Tata, Phys. Rev. D 53, 6241 (1996) [arXiv:hep-ph/9512383]
 ; H. Baer, C. h. Chen, M. Drees, F. Paige and X. Tata, Phys. Rev. D 58, 075008 (1998) [arXiv:hep-ph/9802441];
 S. Abdullin et al. [CMS Collaboration], J. Phys. G 28, 469 (2002) [arXiv:hep-ph/9806366].
- [88] D. N. Spergel et al., Astrophys. J. Suppl. 148 (2003) 175 [arXiv:astro-ph/0302209].
- [89] H. Baer, C. Balazs, A. Belyaev, T. Krupovnickas and X. Tata, JHEP 0306 (2003) 054 [arXiv:hep-ph/0304303].
- [90] B. A. Campbell, D. W. Maybury and B. Murakami, JHEP 0403, 052 (2004) [arXiv:hep-ph/0311244].
- [91] K. Abe et al. [Belle Collaboration], Phys. Rev. Lett. 91, 261602 (2003) [arXiv:hep-ex/0308035];
 B. Aubert et al. [BABAR Collaboration], arXiv:hep-ex/0403026.
- [92] F. Borzumati, C. Greub, T. Hurth and D. Wyler, Phys. Rev. D 62, 075005 (2000) [arXiv:hep-ph/9911245]; M. Ciuchini, G. Degrassi, P. Gambino and G. F. Giudice, Nucl. Phys. B 534, 3 (1998) [arXiv:hep-ph/9806308]; G. Degrassi, P. Gambino and G. F. Giudice, JHEP 0012, 009 (2000) [arXiv:hep-ph/0009337]; M. Carena, D. Garcia, U. Nierste and C. E. M. Wagner, Phys. Lett. B 499, 141 (2001) [arXiv:hep-ph/0010003].
- [93] F. Borzumati, C. Greub and Y. Yamada, arXiv:hep-ph/0305063; F. Borzumati, C. Greub and Y. Yamada, arXiv:hep-ph/0311151.
- [94] Y. Grossman, G. Isidori and M. P. Worah, Phys. Rev. D 58, 057504 (1998) [arXiv:hep-ph/9708305];
 E. Lunghi and D. Wyler, Phys. Lett. B 521, 320 (2001) [arXiv:hep-ph/0109149];
 R. Harnik, D. T. Larson, H. Murayama and A. Pierce, arXiv:hep-ph/0212180;
 S. Khalil and E. Kou, Phys. Rev. D 67, 055009 (2003) [arXiv:hep-ph/0212023];
 G. L. Kane, P. Ko, H. b. Wang,
 - C. Kolda, J. h. Park and L. T. Wang, Phys. Rev. Lett. **90**, 141803 (2003) [arXiv:hep-ph/0304239];
 - D. Chakraverty, E. Gabrielli, K. Huitu and S. Khalil, arXiv:hep-ph/0306076. B. Dutta, C. S. Kim, S. Oh and G. h. Zhu, arXiv:hep-ph/0312389.
- [95] M. Ciuchini, E. Franco, A. Masiero and L. Silvestrini, Phys. Rev. D **67**, 075016 (2003) [arXiv:hep-ph/0212397,v3].
- [96] J. Hisano and Y. Shimizu, Phys. Lett. B 565, 183 (2003) [arXiv:hep-ph/0303071]; Phys. Lett. B 581, 224 (2004) [arXiv:hep-ph/0308255]; J. Hisano, M. Kakizaki, M. Nagai and Y. Shimizu, arXiv:hep-ph/0407169.
- [97] B. Dutta, Y. Mimura and R. N. Mohapatra, arXiv:hep-ph/0406262 and references therein.
- [98] A. Rossi, Phys. Rev. D **66**, 075003 (2002) [arXiv:hep-ph/0207006].
- [99] P. H. Chankowski, J. R. Ellis, S. Pokorski, M. Raidal and K. Turzynski, arXiv:hep-ph/0403180.
- [100] B. Dutta and R. N. Mohapatra, Phys. Rev. D 68, 056006 (2003) [arXiv:hep-ph/0305059].
- [101] A. Ibarra and G. G. Ross, arXiv:hep-ph/0312138;

- [102] A. Ioannisian and A. Pilaftsis, Phys. Rev. D 62, 066001 (2000) [arXiv:hep-ph/9907522]; R. Kitano, Phys. Lett. B 481, 39 (2000) [arXiv:hep-ph/0002279]; S. J. Huber and Q. Shafi, Phys. Lett. B 583, 293 (2004) [arXiv:hep-ph/0309252].
- [103] S. J. Huber, Nucl. Phys. B 666, 269 (2003) [arXiv:hep-ph/0303183].
- [104] D. F. Carvalho, M. E. Gomez and J. C. Romao, Phys. Rev. D **65**, 093013 (2002) [arXiv:hep-ph/0202054].
- [105] F. Deppisch and J. W. F. Valle, arXiv:hep-ph/0406040.
- [106] V. Cirigliano, A. Kurylov, M. J. Ramsey-Musolf and P. Vogel, arXiv:hep-ph/0404233; V. Cirigliano, A. Kurylov, M. J. Ramsey-Musolf and P. Vogel, arXiv:hep-ph/0406199.
- [107] W. Grimus and L. Lavoura, Phys. Rev. D 66, 014016 (2002) [arXiv:hep-ph/0204070].
- [108] See for example, J. G. Korner, A. Pilaftsis and K. Schilcher, Phys. Lett. B **300**, 381 (1993) [arXiv:hep-ph/9301290]; A. Ilakovac and A. Pilaftsis, Nucl. Phys. B **437**, 491 (1995) [arXiv:hep-ph/9403398]; J. I. Illana and T. Riemann, Phys. Rev. D **63**, 053004 (2001) [arXiv:hep-ph/0010193]; J. I. Illana and M. Masip, Phys. Rev. D **67**, 035004 (2003) [arXiv:hep-ph/0207328]; J. Cao, Z. Xiong and J. M. Yang, Eur. Phys. J. C **32** (2004) 245 [arXiv:hep-ph/0307126], and references therein.
- [109] See for example, J. Hisano, M. M. Nojiri, Y. Shimizu and M. Tanaka, Phys. Rev. D 60, 055008 (1999) [arXiv:hep-ph/9808410]; F. Deppisch, H. Pas, A. Redelbach, R. Ruckl and Y. Shimizu, Phys. Rev. D 69, 054014 (2004), and references therein.
- [110] D. Binosi and L. Theussl, Comput. Phys. Commun. 161, 76 (2004) [arXiv:hep-ph/0309015].